## The Collapsing Universe

Understanding "accelerating redshifts" using physics from the 1920's

Bill Sumner

Ellensburg High School June 4, 2015













# Nebula

#### A RELATION BETWEEN DISTANCE AND RADIAL VELOCITY AMONG EXTRA-GALACTIC NEBULAE

By Edwin Hubble

MOUNT WILSON OBSERVATORY, CARNEGIE INSTITUTION OF WASHINGTON

Communicated January 17, 1929







## (Curvature of geometry) = (Mass)

















## **Examples of High-Redshift SNe**

HST observations of SNe in distant galaxies (*Riess et al.*)

Note: you need to ...

- Detect them
- Measure the light curves
- Do the K-corrections
- Get the redshifts







#### Three Americans Share 2011 Nobel Prize in Physics

Published October 04, 2011 / Associated Press



The Royal Swedish Academy of Sciences says American Saul Perlmutter, U.S.-Australian citizen Brian Schmidt and U.S. scientist Adam Riess (pictured here in 2008) share the 2011 Nobel Prize in physics. The trio were honored Tuesday, Oct. 4, 2011 "for the discovery of the acclerating expansion of the universe through observations of distant supernovae." (AP PHOTO/THE JOHN D. AND CATHERINE T. MACARTHUR FOUNDATION, GAIL BURTON)

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STOCKHOLM – Three U.S.-born scientists won the Nobel <u>Prize</u> in physics Tuesday for discovering that the universe is expanding at an accelerating pace, a stunning revelation that suggests the cosmos will eventually freeze to ice.



## (Curvature of geometry) = (Mass)

**+** X

#### **Contents of the Universe: Summary**



•  $\Omega_0 = 1.00 \pm 0.02$ 

• 
$$\Omega_m \approx 0.27 \pm 20\%$$

 $-\Omega_b\approx 0.045\pm 10\%$ 

 $\circ$  Includes  $\Omega_{visible} \approx 0.005$ 

$$-\Omega_{non-b} \approx 0.22$$

 $\circ$  Includes  $\Omega_v < 0.005$ 

$$-\Omega_{CMBR} \approx 0.0001$$

• 
$$\Omega_{de} \approx 0.73 \pm 10\%$$

• The physical nature of the DE is currently completely unknown

#### **Matter Dominated Model**



















## (Curvature of geometry) = (Mass)



 $= \frac{1}{4\pi\varepsilon} \frac{|q_1 q_2|}{r^2}$  $|\mathbf{F}|$ 









"In an expanding space all momenta decrease . . . This simple law has an even simpler interpretation in wave mechanics: <u>all wavelengths</u>, being inversely proportional to the momenta, simply <u>expand with space</u>."

**Erwin Schrödinger** 



























NASA and ESA

STScI-PRC12-48a

![](_page_49_Figure_0.jpeg)

# Which interpretation is "true"?

Ignoring: Is either interpretation "true"?

![](_page_51_Figure_0.jpeg)

#### **Contents of the Universe: Summary**

![](_page_52_Figure_1.jpeg)

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![](_page_53_Picture_0.jpeg)

![](_page_54_Figure_0.jpeg)

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![](_page_56_Figure_0.jpeg)

![](_page_57_Figure_0.jpeg)

![](_page_58_Picture_0.jpeg)

![](_page_59_Picture_0.jpeg)

katoon.org/ehs/

#### HUBBLE REDSHIFT REVISITED

#### W.Q.SUMNER

Abstract. Vacuum permittivity is the scalar in Maxwell's equations that determines the speed of light and the strength of electrical fields. Einstein proved that vacuum permittivity changes with spacetime curvature. In a Friedmann universe, spacetime curvature changes with time, shifting both photons and atomic emissions. A photon today has a different wavelength than it did yesterday. Yesterday, an atom emitted a photon with a different wavelength than it emits today. To understand Hubble redshift in a Friedmann universe, both these changes must be taken into account. When they are, Hubble red- shift implies that the Friedmann universe is closed and collapsing. During collapse, both atomic emissions and photons blueshift. Atomic emissions blueshift about twice as much as photons do. Blueshifted photons seen in a telescope are redder than reference photons emitted today in the observatory. With this insight, supernovae redshift observations are beautifully fit simply using the physics of Maxwell, Einstein, Bohr, and Friedmann from the 1920's. There is no need to postulate dark energy. Supernovae redshift data imply that the universe is very nearly flat and will collapse in about 9.6 billion years. High-z redshift observations up to 11.9 suggest that the universe is at least 2000 billion years old. This is more than a hundred times greater than a typical star's lifetime. This makes it likely that most dark matter is the residue of stellar evolution.